

Other curvature tensors

Ricci

$$R_{\alpha\beta} = R^{\gamma}_{\alpha\gamma\beta}$$

10 indep. components

Symmetric

$$\begin{aligned} R_{\alpha\beta} &= g^{\gamma\delta} R_{\gamma\alpha\delta\beta} \\ &= g^{\gamma\delta} R_{\delta\beta\gamma\alpha} \\ &= R_{\beta\alpha} \end{aligned}$$

Scalar Curvature

$$R = R^{\alpha}_{\alpha}$$

Einstein Curvature

$$G_{\alpha\beta} = R_{\alpha\beta} - \frac{1}{2} g_{\alpha\beta} R$$

Symmetric

Note: $G_{\alpha\beta};^{\beta} = 0$

Contracted Bianchi identity

Why?

$$\text{Bianchi: } R^{\alpha}_{\beta\gamma\delta;\epsilon} + R^{\alpha}_{\beta\delta\epsilon;\gamma} + R^{\alpha}_{\beta\epsilon\gamma;\delta} = 0$$

$$\text{Contract on } \alpha\gamma \Rightarrow R_{\beta\delta;\epsilon} + R^{\alpha}_{\beta\delta\epsilon;\alpha} - R_{\beta\epsilon;\delta} = 0$$

$$\text{Contract on } \beta\delta \Rightarrow R_{;\epsilon} - R^{\alpha}_{\epsilon;\alpha} - R^{\beta}_{\epsilon;\beta} = 0$$

$$\Rightarrow R_{;E} - 2R_{E^{\alpha};\alpha} = 0$$

$$\begin{aligned} \text{But } G_{\alpha}{}^{\beta}{}_{;E} &= R_{\alpha}{}^{\beta}{}_{;E} - \frac{1}{2}(g_{\alpha}{}^{\beta}R)_{;E} \\ &= R_{\alpha}{}^{\beta}{}_{;E} - \frac{1}{2}R_{;E} = 0 \end{aligned}$$

Weyl Curvature

$$C^{\alpha\beta}{}_{\gamma\delta} = R^{\alpha\beta}{}_{\gamma\delta} - 2\delta^{\alpha}{}_{\gamma}\delta^{\beta}{}_{\delta}R + \frac{1}{3}R\delta^{\alpha}{}_{\gamma}\delta^{\beta}{}_{\delta}$$

Same symmetries as $R_{\alpha\beta\gamma\delta}$

but contract any 2 indices $\rightarrow 0$.

Riemann can be decomposed into Ricci + Weyl

"Trace"
Part

"Trace Free"
Part

10 indep
components

- Weyl vanishes in $\dim < 4$

- IF $C^{\alpha\beta}{}_{\gamma\delta} = 0 \Leftrightarrow$ "conformally flat"

$\Leftrightarrow$ Can Find coords such that

$$ds^2 = e^{2\phi} (\eta_{\mu\nu} dx^{\mu} dx^{\nu})$$

$$\phi = \phi(x^{\alpha})$$

Flat Space $\rightarrow$ Curved Space

Einstein Equivalence principle: in every LLF, all nongravitational physics is that of SR.

How to implement: ("comma $\rightarrow$ semicolon rule")

If you write a physical law in geometric form in SR,

The same law applies in GR, except $\partial_\mu \rightarrow \nabla_\mu$
or $A_{,\mu} \rightarrow A_{;\mu}$

Example: in SR, $\nabla \cdot \tilde{T} = 0$ meaning $T^{\alpha\beta}_{;\beta} = 0$

$\tilde{T}$ = stress-energy tensor.

- so in LLF, at origin $T^{\hat{\alpha}\hat{\beta}}_{;\hat{\beta}} = 0$ (E.P.)

- but in LLF at origin $T^{\hat{\alpha}\hat{\beta}}_{;\hat{\beta}} = T^{\hat{\alpha}\hat{\beta}}_{;\hat{\beta}} = 0$

- Geometric law, true in LLF $\Rightarrow$ True in all frames

$$\Rightarrow \underline{T^{\alpha\beta}_{;\beta} = \nabla \cdot \tilde{T} = 0}$$

Similarly, Maxwell: $F^{\alpha\beta}_{;\beta} = 4\pi J^\alpha \Rightarrow F^{\alpha\beta}_{;\beta} = 4\pi J^\alpha$

$$F_{[\alpha\beta];\gamma} = 0 \Rightarrow F_{[\alpha\beta;\gamma]} = 0$$

$$m^\alpha = F^{\alpha\beta} g_{\beta\mu} u^\mu \Rightarrow m^\alpha = F^{\alpha\beta} g_{\beta\mu} u^\mu$$

Ambiguities [curvature coupling]

Example

In SR, consider Maxwell: $F_{[\alpha\beta],\gamma} = 0$ $F^{\alpha\beta}_{;\beta} = 4\pi J^\alpha$

define 4-potential A^α so $F_{\alpha\beta} = A_{\beta,\alpha} - A_{\alpha,\beta}$

$\Rightarrow F_{[\alpha\beta],\gamma} = 0$ automatically

also $F^{\alpha\beta}_{;\beta} = 4\pi J^\alpha$

$$\Rightarrow A^{\beta}_{;\alpha\beta} - A_{\alpha;\beta}{}^{\beta} = 4\pi J_{\alpha} \quad (1)$$

In SR This is same as $A^{\beta}_{;\beta\alpha} - A_{\alpha;\beta}{}^{\beta} = 4\pi J_{\alpha}$ (2)

SR $\Rightarrow$ GR

either $A^{\beta}_{;\alpha\beta} - A_{\alpha;\beta}{}^{\beta} = 4\pi J_{\alpha} \quad (1')$

or $A^{\beta}_{;\beta\alpha} - A_{\alpha;\beta}{}^{\beta} = 4\pi J_{\alpha} \quad (2')$

(1') is same as $A^{\beta}_{;\beta\alpha} - A_{\alpha;\beta}{}^{\beta} + R_{\alpha\beta} A^{\beta} = 4\pi J_{\alpha}$

Which is right? (1') or (2')?

Here go back to basic law $F_{[\alpha\beta],\gamma} = 0 \Rightarrow F_{[\alpha\beta;\gamma]} = 0$

$$F^{\alpha\beta}_{;\beta} = 4\pi J^\alpha \Rightarrow F^{\alpha\beta}_{;\beta} = 4\pi J^\alpha$$

so define A_α s.t. $F_{\alpha\beta} = A_{\beta;\alpha} - A_{\alpha;\beta}$

Then (1') is correct.

Curvature coupling occurs when:

① There are 2nd deriv + order matters

② Can't treat systems as localized in LLF
ie. extended objects

⇒ need physical reason

example: HW problem
precession of equinoxes,

GR in 2 sentences:

- ① - Spacetime curvature tells matter how to move.
- ② - Matter tells spacetime how to curve.

We have seen ①:

all determined
by metric $g_{\mu\nu}$

Freely falling particles move on
geodesics $\nabla_{\vec{u}} \vec{u} = 0$

Geodesic deviation governed by
curvature $\frac{D^2 \vec{n}}{d\tau^2} = \tilde{\tilde{R}}(-, \vec{u}, \vec{u}, \vec{n})$

Now For ②.

We want curvature = mass-energy.
an equation

Want a tensor equation.

Tensor expression for mass-energy is
 $\tilde{\tilde{T}}$, stress-energy tensor.

So we want $\tilde{\tilde{A}} = K \tilde{\tilde{T}}$
something to do with curvature $\uparrow$ constant $\uparrow$

$$\text{or } A_{\alpha\beta} = K T_{\alpha\beta}$$

$\tilde{A}$ should satisfy

- ① Rank 2 + symmetric (because $\tilde{T}$ is)
- ② Built From Riemann + metric
- ③ Linear in Riemann (it's a curvature!)
- ④ Vanishes in Flat space
- ⑤ $A^{\alpha\beta}_{;\beta} = 0$ (Because $T^{\alpha\beta}_{;\beta} = 0$)
Conservation of mass-energy

Only 1 tensor satisfies the above:

Einstein curvature tensor $\tilde{G} \Rightarrow \boxed{G_{\mu\nu} = K T_{\mu\nu}}$

PDEs
that determine
metric $g_{\mu\nu}$

Important
Note on point ⑤.

$G^{\alpha\beta}_{;\beta} = 0$ automatically
(Bianchi identity)

$G^{\alpha\beta}_{;\beta}$ is not a consequence of $T^{\alpha\beta}_{;\beta}$

IF it were, Then the 10 components of $G_{\mu\nu} = K T_{\mu\nu}$
would determine all 10 components of $g_{\mu\nu}$.

But actually, $G^{\alpha\beta}_{;\beta} = 0$ and $T^{\alpha\beta}_{;\beta} = 0$ independently
 $\uparrow$
4 equations

So $G_{\alpha\beta} = K T_{\alpha\beta}$ has only 6 degrees of Freedom ($\begin{matrix} 10 \text{ eqs} \\ - 4 \text{ constraints} \end{matrix}$)
 $\Rightarrow$ only 6 components of $g_{\alpha\beta}$ determined.
Other 4 components represent Freedom in choosing coordinates.