

Physics 236

website

www.tapir.caltech.edu/~vscheel/P236Overview:

Introduce math machinery of general relativity using

- Special Relativity
- E+M
- Physics in curved spacetime

Einstein's Theory of Gravitation: GR

Special topics

- Exact solutions
- Gravitational waves
- Cosmology
- Black holes
- Numerical Relativity

Special Relativity Revisited

→ coordinates are just labels!

spacetime: Set of events. (t, x, y, z) labels events.

reference frame: coordinate system

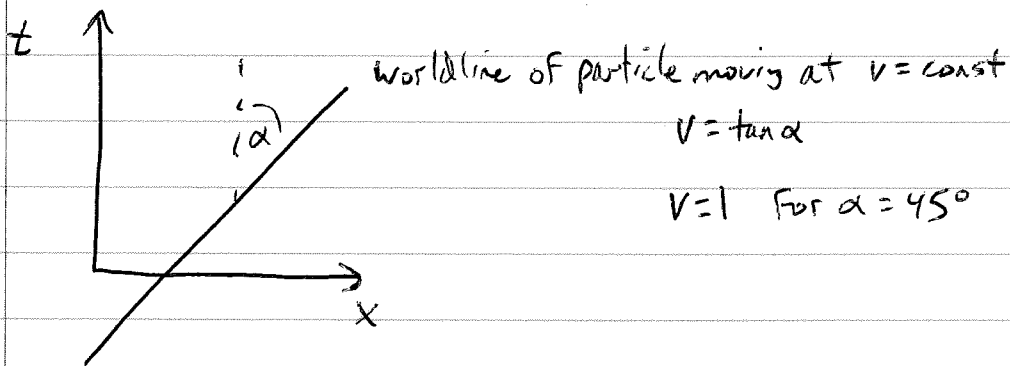
inertial reference frame (Lorentz frame): Particles w/ no force move in straight line.

- I. Relativity Principle: cannot distinguish inertial frames
- II. There is a universal speed limit c

Convention $c \equiv 1$ measure time in cm

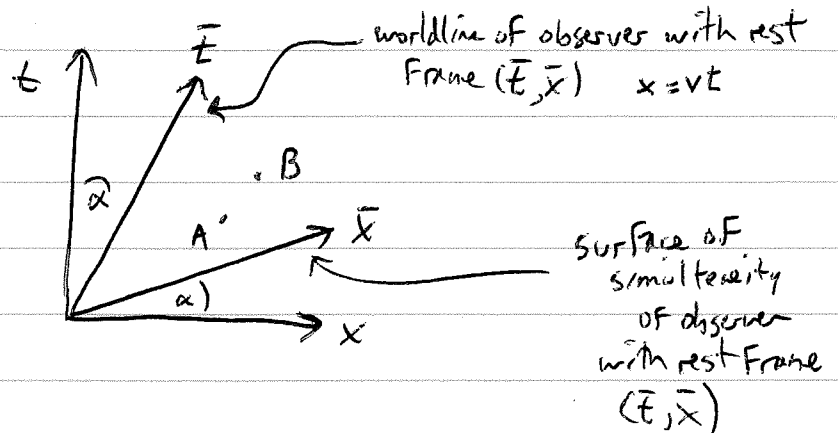
I + II $\Rightarrow$ All of SR

Spacetime Diagrams



Other observers

Frame $(\bar{t}, \bar{x})$ moves with velocity $v = \tan \alpha$ w.r.t. Frame (t, x)



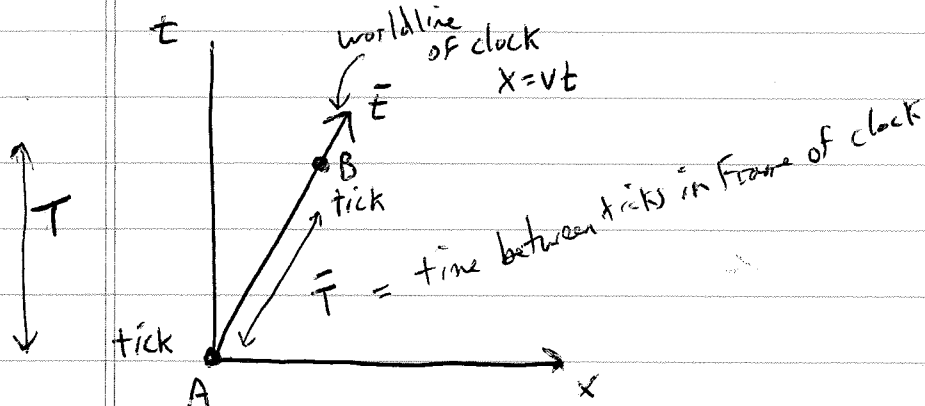
Spacetime interval between 2 events $A+B$ (arbitrary)

$$\bar{t} = vx$$

$$\Delta S^2 \equiv -(\Delta t)^2 + (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2 = -(\Delta \bar{t})^2 + (\Delta \bar{x})^2 + (\Delta \bar{y})^2 + (\Delta \bar{z})^2$$

independent of Frame i.e. invariant

Example: time dilation



interval from A to B :

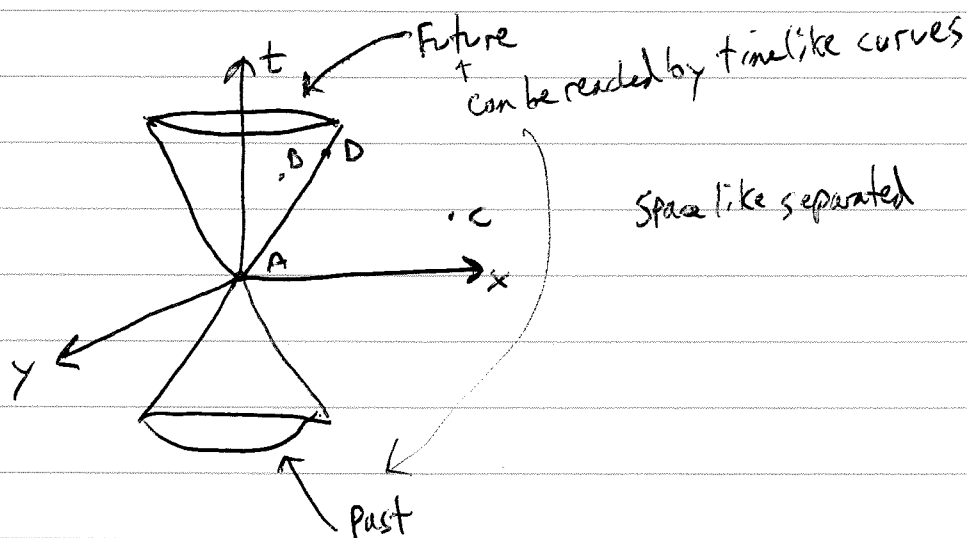
$$\Delta s^2 = -\bar{t}^2 + \bar{x}^2 = -t^2 + x^2$$

$$= -\bar{T}^2 = -t^2 + v^2 t^2$$

$$= -T^2(1-v^2)$$

$$\Rightarrow \boxed{T = \gamma \bar{T}}$$

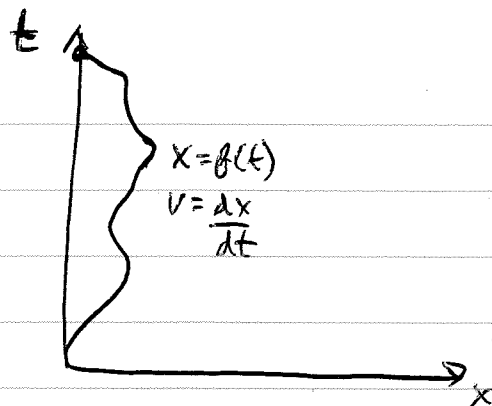
light cone



A vs. C.	$\Delta s^2 > 0$	:	spacelike	Δs = "proper separation"
A vs. B	$\Delta s^2 < 0$	:	timelike	$\sqrt{-\Delta s^2}$ = "proper time" $\equiv \Delta \tau$
A vs. D	$\Delta s^2 = 0$	:	null	

Arbitrary timelike curve

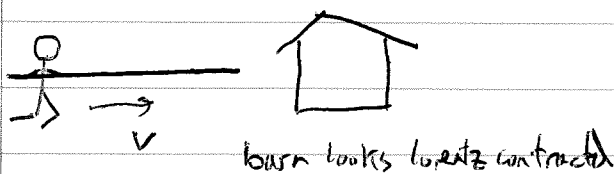
$$\begin{aligned} \text{Proper time} &= \int \sqrt{-ds^2} \\ &= \int \sqrt{1 - v^2} dt \\ &= \int \sqrt{1 - v^2} dt \quad \text{less than } \Delta t \end{aligned}$$



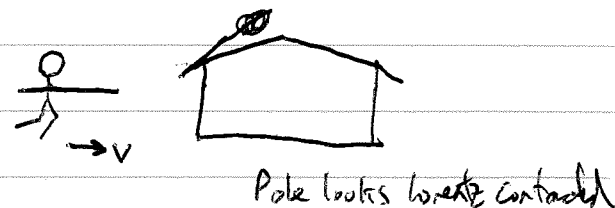
$\Rightarrow$ timelike straight lines are longest

Problem: Pole-vaulter thru barn "paradox"

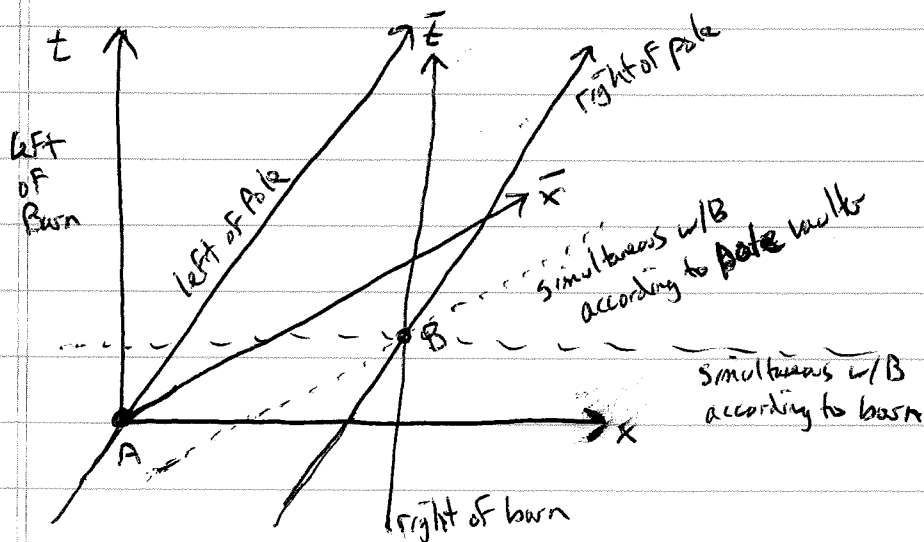
vaulter Frame:



barn Frame:



can Farmer close the barn door on pole?



A = left of pole enters barn
B = right of pole exits barn

- no paradox -
~~Relativity~~
simultaneity is relative

Notation

Label coordinates of some event by X^μ $\mu = 0, 1, 2, 3$

$$X^0 = t$$

$$X^1 = x$$

$$X^2 = y$$

$$X^3 = z$$

Label coordinates $\swarrow$ of same event in different frame by $X^{\bar{\mu}}$

$\bar{\mu} = 0, 1, 2, 3$

$$X^{\bar{0}} = \bar{t}$$

$$X^{\bar{1}} = \bar{x}$$

$$X^{\bar{2}} = \bar{y}$$

$$X^{\bar{3}} = \bar{z}$$

indices are up

Then write interval between X^μ and origin as

$$\Delta s^2 = \sum_{\mu=0}^3 \sum_{\nu=0}^3 \eta_{\mu\nu} X^\mu X^\nu = \eta_{\mu\nu} X^\mu X^\nu$$

$$\eta_{\mu\nu} = \text{"minkowski metric"} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Einstein

Summation

Convention:

repeated up/down

indices are summed

Now transform between coordinates:

$$X^{\bar{\nu}} = \sum_{\mu} \Lambda^{\bar{\nu}}_{\mu} X^{\mu}$$

(summation over μ implied)

$$\Lambda^{\bar{\nu}}_{\mu} = \frac{\partial X^{\bar{\nu}}}{\partial X^{\mu}}$$

Note Some indices are up, others are down. Important

⑦

For SR, consider only subset of transformations
that keep ΔS^2 invariant

$$\begin{aligned}\Delta S^2 &= \eta_{\mu\nu} X^\mu X^\nu \\ &= \eta_{\bar{\mu}\bar{\nu}} X^{\bar{\mu}} X^{\bar{\nu}} = \eta_{\bar{\mu}\bar{\nu}} \Lambda^{\bar{\mu}}_{\nu} X^{\nu} \Lambda^{\bar{\nu}}_{\mu} X^{\mu} \\ &= \underbrace{(\eta_{\bar{\mu}\bar{\nu}} \Lambda^{\bar{\mu}}_{\nu} \Lambda^{\bar{\nu}}_{\mu})}_{\text{must equal } \eta_{\mu\nu}} X^{\mu} X^{\nu}\end{aligned}$$

$\Lambda^{\bar{\mu}}_{\nu}$ that satisfy are Lorentz transformations

• Rotations

• Boosts

$\Lambda^{\bar{\mu}}_{\nu}$ (boost in x direction)

$$= \begin{pmatrix} \gamma & -\gamma v & 0 & 0 \\ -\gamma v & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$X^{\bar{\mu}}$ Frame moving w.r.t.

X^{μ} frame with velocity $+v$

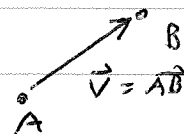
• [Also time reversals, space inversions, but usually we demand Rt. handed coords and time time]

• Poincaré Transformations add translations

Vectors

2 definitions

① Displacement between two events



② Tangent to a curve

curve $C(\lambda)$

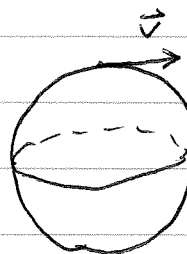


$$\vec{V} = \left. \frac{dC}{d\lambda} \right|_{\lambda=\lambda_0}$$

In Flat space, ① + ② make sense

In curved space, only ② makes sense.

② Vector does not need to lie in
the same space as original manifold.



② $\Rightarrow$ defined locally (i.e. at one point)

Basis Vectors

$$\vec{e}_0, \vec{e}_1, \vec{e}_2, \vec{e}_3$$

example

$$\begin{aligned}\vec{e}_0 &= \hat{t} \\ \vec{e}_1 &= \hat{x} \\ \vec{e}_2 &= \hat{y} \\ \vec{e}_3 &= \hat{z}\end{aligned}$$

Notation: write basis vectors as $\vec{e}_\alpha$ where $\alpha = 0, 1, 2, 3$

For $\vec{e}_\alpha$ to be a basis, any vector $\vec{V}$ can be written

$$\vec{V} = V^0 \vec{e}_0 + V^1 \vec{e}_1 + V^2 \vec{e}_2 + V^3 \vec{e}_3 \equiv \sum_{\alpha} V^{\alpha} \vec{e}_{\alpha} \equiv V^{\alpha} \vec{e}_{\alpha}$$

Einstein Summation
Convention

Note: V^α ($\alpha = 0, 1, 2, 3$) are just numbers
= components of vectors

up, down indices are significant,

Note: "dummy" indices: $V^\alpha \vec{e}_\alpha$ and $V^\beta \vec{e}_\beta$ are the same thing!
and $V^\mu \vec{e}_\mu$ (they are $\vec{V}$)

Note: we usually use $\vec{e}_0 = \hat{t}$, $\vec{e}_1 = \hat{x}$ etc. in SR.

But in general, basis vectors don't need to be orthogonal or unit.

- Vectors can be added or multiplied by a scalar:

$$\vec{V} = a\vec{u} + b\vec{w}$$

$$\text{components are } V^\alpha = a u^\alpha + b w^\alpha$$

- Transformations of vectors into different frames

$\vec{\Delta x}$ displacement vector

$$= (\Delta x)^\mu \vec{e}_\mu = (\Delta x)^{\bar{\mu}} \vec{e}_{\bar{\mu}}$$

$$= (\Delta x)^\mu \Lambda^{\bar{\mu}}_\mu \vec{e}_{\bar{\mu}}$$

$$\Rightarrow \vec{e}_\mu = \Lambda^{\bar{\mu}}_\mu \vec{e}_{\bar{\mu}}$$

(only makes sense in flat space)

$$\text{Components: } \vec{V} = V^\mu \vec{e}_\mu = V^{\bar{\mu}} \vec{e}_{\bar{\mu}} = V^{\bar{\mu}} \Lambda^{\mu}_{\bar{\mu}} \vec{e}_\mu$$

$$\Rightarrow V^\mu = \Lambda^{\mu}_{\bar{\mu}} V^{\bar{\mu}}$$