Solutions Ph 236a – Week 3

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Problem 1

Part (a)

We are given $u \wedge v = u \otimes v - v \otimes u$ so we can see that

$$v \wedge u = v \otimes u - u \otimes v = -(u \otimes v - v \otimes u) = -u \wedge v \tag{1.1}$$

so the wedge product is antisymmetric. When looking at p- and q-forms $\Omega_1 = \frac{1}{p!} \Omega_{\alpha\beta...} u_1 \wedge ... \wedge u_p$ and $\Omega_2 = \frac{1}{q!} \omega_{\alpha\beta...} v_1 \wedge ... \wedge v_2$ and so the wedge product will produce an antisymmetric (p+q)-form

$$\boldsymbol{\Omega_1} \wedge \boldsymbol{\Omega_2} = \left(\frac{1}{p!} \Omega_{\alpha\beta\dots} (u_1 \wedge \dots \wedge u_p)\right) \wedge \left(\frac{1}{q!} \omega_{\mu\nu\dots} (v_1 \wedge \dots \wedge v_q)\right) \\
= \left(\frac{1}{p!q!} \Omega_{\alpha\beta\dots} \omega_{\mu\nu\dots}\right) u_1 \wedge \dots \wedge u_p \wedge v_1 \wedge \dots \wedge v_q$$
(1.2)

Similarly,

$$\boldsymbol{\Omega_2} \wedge \boldsymbol{\Omega_1} = \left(\frac{1}{p!} \Omega_{\alpha\beta\dots} (v_1 \wedge \dots \wedge v_q)\right) \wedge \left(\frac{1}{q!} \omega_{\mu\nu\dots} (u_1 \wedge \dots \wedge u_p)\right)$$
$$= \left(\frac{1}{p!q!} \Omega_{\alpha\beta\dots} \omega_{\mu\nu\dots}\right) v_1 \wedge \dots \wedge v_q \wedge u_1 \wedge \dots \wedge u_p \tag{1.3}$$

so we can see that this equation has all of the same terms as $\Omega_1 \wedge \Omega_2$ up to a possible minus sign. In order to determine the sign, use the anticommutator relation for the wedge product shown above.

$$u_{1} \wedge \dots \wedge u_{p} \wedge v_{1} \wedge \dots \wedge v_{q}$$

$$= (-1)u_{1} \wedge \dots \wedge u_{p-1} \wedge v_{1} \wedge u_{p} \wedge v_{2} \dots \wedge v_{q}$$

$$= (-1)^{p}v_{1} \wedge u_{1} \wedge \dots \wedge u_{p} \wedge v_{2} \wedge \dots \wedge v_{q}$$

$$= \dots (\text{repeat this process for each } v_{i}) \dots$$

$$= (-1)^{pq}v_{1} \wedge \dots \wedge v_{q} \wedge u_{1} \wedge \dots \wedge u_{p} \qquad (1.4)$$

which is the same as the term as in $\Omega_2 \wedge \Omega_1$ up to $(-1)^{pq}$, so because the coefficients are all equivalent, then we can see that $\Omega_1 \wedge \Omega_2 = (-1)^{pq} \Omega_2 \wedge \Omega_1$

Part (b)

By definition of the exterior derivative, if Ω_1 is a *p*-form, then $d\Omega_1$ is a (p+1)-form. To get the desired result, start with equation (3) from the homework and

use the commutation relation derived above.

$$\tilde{d}(\boldsymbol{\Omega}_{1} \wedge \boldsymbol{\Omega}_{2}) = \tilde{d}\boldsymbol{\Omega}_{1} \wedge \boldsymbol{\Omega}_{2} + (-1)^{p}\boldsymbol{\Omega}_{1} \wedge \tilde{d}\boldsymbol{\Omega}_{2}
\Rightarrow \tilde{d}((-1)^{pq}\boldsymbol{\Omega}_{2} \wedge \boldsymbol{\Omega}_{1}) = (-1)^{(p+1)q}\boldsymbol{\Omega}_{2} \wedge \tilde{d}\boldsymbol{\Omega}_{1} + (-1)^{p}(-1)^{p(q+1)}\tilde{d}\boldsymbol{\Omega}_{2} \wedge \boldsymbol{\Omega}_{1}
\Rightarrow \tilde{d}(\boldsymbol{\Omega}_{2} \wedge \boldsymbol{\Omega}_{1}) = (-1)^{q}\boldsymbol{\Omega}_{2} \wedge \tilde{d}\boldsymbol{\Omega}_{1} + (-1)^{2p}\tilde{d}\boldsymbol{\Omega}_{2} \wedge \boldsymbol{\Omega}_{1}
\Rightarrow \tilde{d}(\boldsymbol{\Omega}_{2} \wedge \boldsymbol{\Omega}_{1}) = \tilde{d}\boldsymbol{\Omega}_{2} \wedge \boldsymbol{\Omega}_{1} + (-1)^{q}\boldsymbol{\Omega}_{2} \wedge \tilde{d}\boldsymbol{\Omega}_{1}$$
(1.5)

Which is the result we are trying to prove.

Problem 2

Part (a)

Recall that we can always find a frame that locally looks like flat spacetime. In that frame, the Faraday tensor is

$$F_{\mu\nu} = \begin{bmatrix} 0 & -E^1 & -E^2 & -E^3 \\ E^1 & 0 & B^3 & -B^2 \\ E^2 & -B^3 & 0 & B^1 \\ E^3 & B^2 & -B^1 & 0 \end{bmatrix},$$

$$F^{\mu\nu} = \begin{bmatrix} 0 & E^1 & E^2 & E^3 \\ -E^1 & 0 & B^3 & -B^2 \\ -E^2 & -B^3 & 0 & B^1 \\ -E^3 & B^2 & -B^1 & 0 \end{bmatrix}.$$
 (2.1)

Since F is a tensor, if we contract its indices, we get something that is invariant. For example, $F_{\mu\nu}F^{\mu\nu}$ is invariant. Note that this operation corresponds to multiplying the matrices $F_{\mu\nu}$ and $F^{\mu\nu}$ element-wise and then adding up all entries. We find that

$$F_{\mu\nu}F^{\mu\nu} = -2\|\underline{\tilde{E}}\|^2 + 2\|\underline{\tilde{B}}\|^2 = 2(B^2 - E^2).$$
(2.2)

Since $2(B^2 - E^2)$ is invariant, it follows that $B^2 - E^2$ is also invariant.

Recall that $F_{\mu\nu}$ is antisymmetric, thus its dual ${}^*F_{\alpha\beta}$ exists and is defined as

$${}^{*}F_{\alpha\beta} = \frac{1}{2!}F^{\mu\nu}\epsilon_{\mu\nu\alpha\beta} = \frac{1}{2} \begin{bmatrix} 0 & 2B^{1} & -2(-B^{2}) & 2B^{3} \\ -2B^{1} & 0 & 2E^{3} & -2E^{2} \\ 2(-B^{2}) & -2E^{3} & 0 & 2E^{1} \\ -2B^{3} & 2E^{2} & -2E^{1} & 0 \end{bmatrix} = \\ = \begin{bmatrix} 0 & B^{1} & B^{2} & B^{3} \\ -B^{1} & 0 & E^{3} & -E^{2} \\ -B^{2} & -E^{3} & 0 & E^{1} \\ -B^{3} & E^{2} & -E^{1} & 0 \end{bmatrix}.$$
(2.3)

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Since ${}^*F_{\alpha\beta}$ is a tensor, the contraction ${}^*F_{\alpha\beta}F^{\alpha\beta}$ is invariant, hence

$${}^{*}F_{\alpha\beta}F^{\alpha\beta} = 4\underline{E}\cdot\underline{B} \tag{2.4}$$

is also invariant. Thus $E \cdot B$ is invariant.

Part (b)

If a quantity is a spacetime invariant, it also must be an invariant under spatial rotations, i.e. a spatial scalar. In a given Lorentz frame, there are only three ways to make a spatial scalar from \underline{E} and \underline{B} : these are $\underline{E} \cdot \underline{E}$, $\underline{B} \cdot \underline{B}$, and $\underline{B} \cdot \underline{E}$ (and combinations of these scalars).

Now suppose that there were three independent *spacetime* invariants that can be made from $\underline{E} \cdot \underline{E}$, $\underline{B} \cdot \underline{B}$, and $\underline{B} \cdot \underline{E}$. This would imply that each of these spatial scalars is actually a spacetime scalar. However, it is easy to verify (e.g. by carrying out an explicit Lorentz transformation) that $\underline{E} \cdot \underline{E}$ and $\underline{B} \cdot \underline{B}$ are not spacetime scalars.

Alternatively, one can show that all combinations of $F^{\mu\nu}$ and ${}^*F^{\mu\nu}$, where all indices are contracted, involve only $B^2 - E^2$ or $E \cdot B$. This is more difficult.

Since both tensors are antisymmetric we have

$$F^{\mu}_{\ \mu} = {}^{*}F^{\mu}_{\ \mu} = 0, \tag{2.5}$$

which is not a combination of \underline{E} and \underline{B} . Thus we need to consider expressions with two tensors. The possible combinations are

$$F_{\mu\nu}F^{\mu\nu} = -F_{\mu\nu}F^{\nu\mu} = 2(B^2 - E^2),$$

$${}^*F_{\mu\nu}{}^*F^{\mu\nu} = -{}^*F_{\mu\nu}{}^*F^{\nu\mu} = 2(E^2 - B^2),$$

$$F_{\mu\nu}{}^*F^{\mu\nu} = -F_{\mu\nu}{}^*F^{\nu\mu} = 4\underline{E}\cdot\underline{B},$$
(2.6)

which are just multiples of $B^2 - E^2$ or $\underline{E} \cdot \underline{B}$. Note that inserting a Levi-Civita tensor would simply convert one of the tensors to its dual and it would not give a different combination of $F^{\mu\nu}$ and $*F^{\mu\nu}$ that what was considered above.

Unfortunately, the above is not quite enough to show that the only invariants are $B^2 - E^2$ and $\tilde{E} \cdot \tilde{B}$. One can also construct scalars with more than just two tensors. For example,

$$F^{\mu}_{\ \nu} F^{\nu}_{\ \lambda} F^{\lambda}_{\ \mu} = (-F_{\nu}^{\ \mu})(-F_{\lambda}^{\ \nu})(-F_{\mu}^{\ \lambda}) = -F^{\nu}_{\ \mu} F^{\lambda}_{\ \nu} F^{\mu}_{\ \lambda}$$
$$= -F^{\nu}_{\ \mu} F^{\mu}_{\ \lambda} F^{\lambda}_{\ \nu} , \qquad (2.7)$$

since we can relabel contracted indices, the above says x = -x and thus $x = F^{\mu}_{\ \nu} F^{\nu}_{\ \lambda} F^{\lambda}_{\ \mu} = 0$, in this particular example. There seems to be no easy argument that one could apply to *any* higher order combination of F and *F with all indices contracted.

Another proof that the only invariants formed by E and B are $B^2 - E^2$ and $E \cdot B$ goes as follows. Consider the vector

$$\underline{F} = \underline{E} + i\underline{B} = (E_x + iB_x, E_y + iB_y, E_z + iB_z), \qquad (2.8)$$

where *i* is the imaginary unit. Now consider a Lorentz boost in the (t, x), plane, i.e. let $y = (v_x, 0, 0)$. Applying the usual Lorentz transformations for E and B, namely,

$$\begin{split} \mathcal{E}'_{\parallel} &= \mathcal{E}_{\parallel} \\ \mathcal{E}'_{\perp} &= \gamma(\mathcal{E}_{\perp} + \mathcal{y} \times \mathcal{B}_{\perp}) \\ \mathcal{B}'_{\parallel} &= \mathcal{B}_{\parallel} \\ \mathcal{B}'_{\perp} &= \gamma(\mathcal{B}_{\perp} - \mathcal{y} \times \mathcal{E}_{\perp}), \end{split}$$
(2.9)

where $E_{\parallel} = (E_x, 0, 0), E_{\perp} = (0, E_y, E_z)$ and similarly for \tilde{B} , we find

$$F'_{x} = E_{x} + iB_{x} = F_{x}$$

$$F'_{y} = \gamma(E_{y} - v_{x}B_{z}) + i\gamma(B_{y} + v_{x}E_{z})$$

$$= \gamma(E_{y} + iB_{y}) + \gamma v_{x}(iE_{z} - B_{z}) = \gamma F_{y} + \gamma v_{x}iF_{z}$$

$$F'_{z} = \gamma(E_{z} + v_{x}B_{y}) + i\gamma(B_{z} - v_{x}E_{y})$$

$$= \gamma(E_{z} + iB_{z}) + \gamma v_{x}(-iE_{y} + B_{y}) = \gamma F_{z} - \gamma v_{x}iF_{y}.$$
(2.10)

We can write the above as

$$F'_{x} = F_{x}$$

$$F'_{y} = \cosh \psi F_{y} + i \sinh \psi F_{z}$$

$$F'_{z} = -i \sinh \psi F_{y} + \cosh \psi F_{z},$$
(2.11)

where $\tanh \psi = v_x$ and so

$$\cosh \psi = \sqrt{\frac{1}{1 - \tanh^2 \psi}} = \sqrt{\frac{1}{1 - v_x^2}} = \gamma$$
$$\sinh \psi = \sqrt{\cosh^2 \psi - 1} = \sqrt{\gamma^2 - 1} = \sqrt{\frac{1}{1 - v_x^2} - 1} = \sqrt{\frac{1 - 1 + v_x^2}{1 - v_x^2}}$$
$$= \sqrt{\frac{v_x^2}{1 - v_x^2}} = \frac{v_x}{\sqrt{1 - v_x^2}} = v_x \gamma.$$
(2.12)

Recall that $\cosh \psi = \cos i \psi$ and $\sinh \psi = -i \sin i \psi$, so we get

$$F'_{x} = F_{x}$$

$$F'_{y} = \cos i\psi F_{y} + \sin i\psi F_{z}$$

$$F'_{z} = -\sin i\psi F_{y} + \cos i\psi F_{z},$$
(2.13)

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and so \tilde{F}' is obtained by rotating \tilde{F} by the imaginary angle $-i\psi$ in 3-space in the (y, z) plane. Similarly, the Lorentz boosts in the (t, y) and (t, z) planes correspond to rotations by imaginary angles in the (x, z) and (x, y) planes in 3-space. The only other Lorentz transformations in 4-space are regular rotations in 3-space, which obviously correspond to rotations by real angles in 3-space.

Thus we have a correspondence between all Lorentz transformations in 4space (rotations in the (t, x), (t, y), (t, z), (x, y), (x, z), and (y, z) planes) to rotations in 3-space by imaginary or real angles. Note that \underline{E} and \underline{B} are independent in \underline{F} and thus by considering all Lorentz transformations of \underline{F} in 4-space, we are actually considering all possible Lorentz transformations of \underline{E} and \underline{B} . All these transformations correspond to rotations of \underline{F} in 3-space (by imaginary or complex angles). But rotations only preserve the length of a vector, thus $\|\underline{F}\|$ is the *only* quantity conserved by rotations and thus it is the *only* quantity conserved by Lorentz transformations of \underline{E} and \underline{B} . Therefore, the only conserved quantity is

$$\|\underline{F}\| = E^2 - B^2 + 2i\underline{E} \cdot \underline{B}, \qquad (2.14)$$

hence the real numbers $E^2 - B^2$ and $E \cdot B$ are the *only* quantities that are independently conserved under Lorentz transformations of E and B.

Also see http://arxiv.org/abs/1309.4185 for a more complicated proof.

Problem 3

Part (a)

We are given that $A_{\alpha\beta}u^{\alpha}u^{\beta} = 0$ for a timelike vector u^{α} . Now let $\vec{u} = \vec{e_t} + \epsilon \vec{e_i}, 0 < \epsilon < 1$ for i = 1, 2, 3. Expanding $A_{\alpha\beta}u^{\alpha}u^{\beta} = 0$, we get

$$A_{00} + \epsilon A_{0i} + \epsilon A_{i0} + \epsilon^2 A_{ii} = 0$$

$$\Rightarrow A_{00} = 0, A_{ii} = 0, A_{i0} = -A_{0i}$$
(3.1)

Now let $\vec{u} = \vec{e}_t + \epsilon \vec{e}_i + \delta \vec{e}_j, 0 < \epsilon, \delta < \frac{1}{2}$ for (i, j) = 1, 2, 3 and $i \neq j$. Then we get

$$A_{00} + \epsilon A_{0i} + \epsilon A_{i0} + \epsilon^2 A_{ii} + \delta A_{0j} + \delta A_{j0} + \epsilon^2 A_{jj} + \epsilon \delta A_{ij} + \delta \epsilon A_{ji} = 0$$

$$\Rightarrow 0 + \epsilon A_{0i} - \epsilon A_{0i} + \epsilon^2 0 + \delta A_{0j} - \delta A_{0j} + \epsilon^2 0 + \epsilon \delta A_{ij} + \delta \epsilon A_{ji} = 0$$

$$\Rightarrow A_{ij} = -A_{ji} \qquad (3.2)$$

thus we have $A_{\alpha\beta} = -A_{\beta\alpha} \forall \alpha, \beta = 0, 1, 2, 3.$

Part (b)

Lets start with the 4-momentum relationship $\vec{p} \cdot \vec{p} = -m^2$ where *m* is the rest mass of an arbitrary particle. Now take the derivative of both sides,

$$\frac{d}{d\tau}(-m^2) = 0 = \frac{d}{d\tau}(\vec{p}\cdot\vec{p}) = 2\vec{p}\cdot\frac{d}{d\tau}\vec{p}$$
(3.3)

However, we know that the Lorentz force is given by the equation,

$$\left(\frac{d}{d\tau}\vec{p}\right)^{\alpha} = qF^{\alpha}{}_{\beta}\,u^{\beta} \tag{3.4}$$

where q is the charge and F is the Faraday tensor. Substituting that back in, we find that

$$0 = p_{\alpha} \left(\frac{d}{d\tau}p\right)^{\alpha} = q\vec{p}_{\alpha}F^{\alpha}_{\ \beta}u^{\beta} = mqF^{\alpha}_{\ \beta}u_{\alpha}u^{\beta}$$
$$\Rightarrow F_{\alpha\beta}u^{\alpha}u^{\beta} = 0$$
(3.5)

Therefore, using the result from Part (a), we can say that F is antisymmetric.

Problem 4

Part (a)

See (2.1) for the Faraday tensor. From it, we can compute

$$F^{\mu}_{\ \nu} = \begin{bmatrix} 0 & E^1 & E^2 & E^3 \\ E^1 & 0 & B^3 & -B^2 \\ E^2 & -B^3 & 0 & B^1 \\ E^3 & B^2 & -B^1 & 0 \end{bmatrix}$$
(4.1)

Now the contraction of F with itself gives us

$$F^{\mu\nu}F_{\mu\nu} = 2(\tilde{B}^2 - \tilde{E}^2)$$
(4.2)

where $\tilde{B}^2 = (B^1)^2 + (B^2)^2 + (B^3)^2$. So then, using equation (7) from the problem set, the energy density is

$$T^{00} = \frac{1}{4\pi} \left(F^{0\mu} F^{0}_{\ \mu} - \frac{2}{4} g^{00} (\tilde{B}^2 - \tilde{E}^2) \right)$$
$$= \frac{1}{4\pi} \left(\tilde{E}^2 + \frac{1}{2} (\tilde{B}^2 - \tilde{E}^2) \right)$$
$$= \frac{1}{8\pi} (\tilde{B}^2 + \tilde{E}^2)$$
(4.3)

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which should be a familiar result. To get the momentum density in general, let us first look at a particular case, the density in the x-direction

$$T^{0x} = \frac{1}{4\pi} \left(F^{0\mu} F^x_{\ \mu} - \frac{1}{2} g^{0x} (\underline{B}^2 - \underline{E}^2) \right) = -\frac{1}{4\pi} (E^y B^z - E^z B^y + 0) \quad (4.4)$$

The momentum density in the i-direction is given by

$$T^{0i} = \frac{1}{4\pi} \left(F^{0\mu} F^{i}_{\mu} \right)$$
$$= \frac{1}{4\pi} \left(\epsilon_{ijk} E^{j} B^{k} \right)$$
(4.5)

Which can be written in 3-vector form as the Poynting vector $\vec{S} = \frac{1}{4\pi}\vec{E} \times \vec{B}$. Finally, for the 3D stress tensor, if i = j, then $g^{ii} = 1$ and otherwise $g^{ij} = 0$ so we can rewrite the 3D spatial metric as $g^{ij} = \delta^{ij}$. To get the rest of the tensor, consider the case where i = x and j = y. Then

$$F^{x\mu}F^y_{\ \mu} = -E^x E^y - B^x B^y$$

and in the case where i = j = x then

$$F^{x\mu}F^{y}_{\ \mu} = -(E^x)^2 + B^2 - (B^x)^2$$

so in general

$$F^{i\mu}F^{j}_{\ \mu} = -E^{i}E^{j} - B^{i}B^{j} + \delta^{ij}B^{2}$$

$$\Rightarrow T^{ij} = \frac{1}{4\pi} \left(-E^{i}E^{j} - B^{i}B^{j} + \frac{1}{2}\delta^{ij}(\bar{B}^{2} + \bar{E}^{2}) \right)$$
(4.6)

Part (b)

Given the general Faraday, we know that part of Maxwell's equations read

$$\partial_{\beta}F^{\alpha\beta} = F^{\alpha\beta}_{\ \beta} = 4\pi J^{\alpha} \tag{4.7}$$

Now let's take the derivative of the Maxwell stress-energy tensor, (noting that $F^{\mu\nu}\partial_{\beta}F_{\mu\nu} = F_{\mu\nu}\partial_{\beta}F^{\mu\nu}$),

$$\partial_{\beta}T^{\alpha\beta} = \frac{1}{4\pi} \left(F^{\beta}_{\ \mu} \partial_{\beta}F^{\alpha\mu} + F^{\alpha\mu}\partial_{\beta}F^{\beta}_{\ \mu} - \frac{1}{4}g^{\alpha\beta}(2F^{\mu\nu}\partial_{\beta}F_{\mu\nu}) \right)$$
$$= \frac{1}{4\pi} \left(-4\pi F^{\alpha\mu}J_{\mu} + \left(F^{\beta}_{\ \mu}\partial_{\beta}F^{\alpha\mu} - \frac{1}{2}F^{\mu\beta}\partial^{\alpha}F_{\mu\beta} \right) \right)$$
(4.8)

Clearly we will have the desired result $\partial_{\beta}T^{\alpha\beta} = -F^{\alpha\beta}J_{\beta}$ if the other terms vanish. We can show that it does by using the antisymmetry of $F^{\mu\nu}$ and the Maxwell equation

$$\partial_{\beta}F_{\alpha\mu} + \partial_{\alpha}F_{\mu\beta} + \partial_{\mu}F_{\beta\alpha} = 0. \tag{4.9}$$

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Using this equation and maniplating upper and lower indices we can show that

$$0^{\alpha} = F^{\beta\mu} \left(\partial_{\beta} F^{\alpha}{}_{\mu} + \partial^{\alpha} F_{\mu\beta} + \partial_{\mu} F_{\beta}{}^{\alpha} \right) = F^{\beta}{}_{\mu} \partial_{\beta} F^{\alpha\mu} - F^{\beta\mu} \partial^{\alpha} F_{\mu\beta} + F^{\mu\beta} \partial_{\beta} F_{\mu}{}^{\alpha} = 2 \left(F^{\beta}{}_{\mu} \partial_{\beta} F^{\alpha\mu} - \frac{1}{2} F^{\beta\mu} \partial^{\alpha} F_{\mu\beta} \right)$$
(4.10)

Thus the extra terms vanish so we get the result $\partial_{\beta}T^{\alpha\beta} = -F^{\alpha\beta}J_{\beta}$.

Part (c)

If the total stress energy tensor is conserved, then $\partial_{\beta} \left(T_{EM}^{\alpha\beta} + T_{matter}^{\alpha\beta} \right) = 0$ and $\partial_{\beta}T_{EM}^{\alpha\beta} = -F^{\alpha\beta}J_{\beta}$ give $\partial_{\beta}T_{matter}^{\alpha\beta} = F^{\alpha\beta}J_{\beta}$. Examine the time and spatial components separately with $\vec{J} = (\rho, \underline{i})$.

First, note that by definition of $T_{matter}^{\alpha\beta}$, T_{matter}^{00} is the energy density and T_{matter}^{0i} is the energy flux in the *i*-th direction. Then we can say

$$\alpha = 0: \partial_t T_{matter}^{00} + \partial_i T_{matter}^{0i} = \frac{d\text{Energy}}{dtdV} = F^{0i} J_i = \underline{j} \cdot \underline{\tilde{E}}$$
(4.11)

which is simply Ohmic heating.

For the spatial components, T_{matter}^{i0} is the density of the *i*-th component of momentum and T_{matter}^{ij} is the flux of the *i*-th component of momentum in the *j*-th direction. Then

$$\alpha = i : \partial_t T^{i0}_{matter} + \partial_i T^{ij}_{matter} = \frac{dp^i}{dtdV} = F^{i0}J_0 + F^{ij}J_j = \rho E^i + \epsilon^{ijk}J_j B_k$$
(4.12)

which is just the Lorentz force equation.

Problem 5

Part (a)

From $T^{\mu\nu}_{,\nu} = 0$ we obtain $0 = T^{\mu\nu}_{,\nu} = T^{\mu0}_{,0} + T^{\mu i}_{,i}$ and so $T^{\mu0}_{,0} = -T^{\mu i}_{,i}$. Recall that $\partial_0 = \partial/\partial t$ and so we find

$$\frac{\partial}{\partial t} \int T^{0\alpha} d^3 x = \int T^{\alpha 0}_{,0} d^3 x = -\int T^{\alpha i}_{,i} d^3 x, \qquad (5.1)$$

where we have assumed that $T^{0\alpha}$ is a smooth function of x^{μ} so that we can interchange differentiation and integration, and we have used that T is symmetric. Applying the divergence theorem to the last integral yields

$$\frac{\partial}{\partial t} \int T^{0\alpha} d^3 x = -\int_S T^{\alpha i} n_i \, dS = 0, \qquad (5.2)$$

where n_i is the normal vector to the surface S. Since $T^{\mu\nu} = 0$ outside some bounded region of space, take the surface S to be a surface that is outside this region and encloses it. Thus on the surface S, $T^{\mu\nu} = 0$ and so the integral over S is zero. Note that the contributions to the volume integral from points outside S is zero, because $T^{\mu\nu}$ vanishes there. Thus our result does indeed apply to the volume integral over all space.

Part (b)

We find

$$\begin{aligned} \frac{\partial^2}{\partial t^2} \int T^{00} x^i x^j \, d^3x &= \frac{\partial}{\partial t} \int T^{00}_{,0} x^i x^j \, d^3x = -\frac{\partial}{\partial t} \int T^{0k}_{,k} x^i x^j \, d^3x \\ &= -\frac{\partial}{\partial t} \left(\int_S T^{0k} x^i x^j n_k \, d^3x - \int T^{0k} (\delta^i_k x^j + x^i \delta^j_k) \right) \\ &= 0 + \frac{\partial}{\partial t} \int (T^{0i} x^j + T^{0j} x^i) \, d^3x, \end{aligned}$$
(5.3)

where we have used integration by parts and the surface term vanishes by the same argument used in part (a). Finally we get

$$\frac{\partial^2}{\partial t^2} \int T^{00} x^i x^j \, d^3 x = \int (T^{0i}_{,0} x^j + T^{0j}_{,0} x^i) \, d^3 x = -\int (T^{ki}_{,k} x^j + T^{kj}_{,k} x^i) \, d^3 x$$
$$= -0 + \int (T^{ki} \delta^j_{\ k} + T^{kj} \delta^i_{\ k}) \, d^3 x = 2 \int T^{ij} \, d^3 x, \qquad (5.4)$$

where we have again used integration by parts and the fact that the surface terms are 0. $\hfill \Box$

Part (c)

We obtain

$$\begin{aligned} \frac{\partial^2}{\partial t^2} \int T^{00} (x^i x_i)^2 \, d^3x &= \frac{\partial}{\partial t} \int T^{00}_{,0} \, (x^i x_i)^2 \, d^3x = -\frac{\partial}{\partial t} \int T^{0k}_{,k} \, (x^i x_i)^2 \, d^3x \\ &= -0 + \frac{\partial}{\partial t} \int T^{0k} 2 (x^j x_j) (\delta^i_k x_i + x^i \delta_{ik}) \, d^3x \\ &= 4 \int T^{k0}_{,0} \, (x^j x_j) x_k \, d^3x = -4 \int T^{kl}_{,l} \, (x^j x_j) x_k \, d^3x \\ &= -0 + 4 \int T^{kl} (\delta^j_l x_j x_k + x^j \delta_{jl} x_k + x^j x_j \delta_{kl}) \, d^3x \\ &= 4 \int (T^{kj} x_j x_k + T^k_j \, x^j x_k + T^k_k \, x^j x_j) \, d^3x \\ &= 8 \int T^{ij} x_i x_j \, d^3x + 4 \int T^k_k \, x^i x_i \, d^3x, \end{aligned}$$
(5.5)

where we have used integration by parts twice.